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Implementation of Two New Spectral Input Options in HOS-Ocean – Part I

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NRL MEMORANDUM REPORT

Implementation of two new spectral input options in HOS-ocean. Part I Mark D. Orzech (NRL 7322) Jie Yu (NRL 7322)

Abstract:

This report describes the implementation and testing of two new options in the higherorder spectral (HOS) model "HOS-ocean", each of which facilitates model initialization with two-dimensional (2D) frequency-directional wave spectra. The first option enables the user to generate a JONSWAP-based 2D spectrum with energy peaks at two different directions, specifying the principal directions, the directional spreads, and the JONSWAP parameters for the individual peaks. The second option facilitates model initialization using a frequency-directional spectral data file. The format of the data file is assumed to be that of the WAVEWATCH III (WW3) 2D spectral files, but the source of the data can be any. A library of bivariate interpolation of scattered data is installed, and used to transfer the input spectral data onto the HOS-ocean model grids in order to initialize the model fields. Results from preliminary test cases are presented and discussed here to evaluate the basic performance of the model utilizing the new features.

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1 INTRODUCTION

This report documents two additions that were made to the public-domain software package HOS-ocean (Ducrozet et al., 2007, 2016), each of which enables the model to initialize wave simulations with a two-dimensional (2D) frequency-directional spectrum. The first case facilitates initialization with a directionally bimodal spectrum (i.e., two wave peaks from different directions). Each peak is specified as a JONSWAP frequency spectrum, to which a cosine-power directional distribution is then applied. The second case facilitates model initialization using a frequency-directional spectrum data file. While the format of the input spectral data is assumed to be that of the output files from WAVEWATCH III (WW3DG, 2019), there is no assumption on the source of the data.

The HOS-Ocean model, developed at Ecole Centrale de Nantes (France), is based on the High-Order Spectral (HOS) method. It can efficiently simulate the propagation of highly nonlinear wave fields in three dimensions (assuming a constant or deep water depth). The HOS method is based on the potential flow theory, and recasts the Laplace Equation and boundary conditions into a system of fully nonlinear evolution equations for the free surface displacement and surface velocity potential. To solve the system, the vertical velocity at the free surface is needed, and obtained through a sequence of evaluations that involve eigenfunction expansion of the surface velocity potential in the wavenumber space (West et al., 1987; Dommermuth and Yue, 1987; among others). The transformation to and from the physical and wavenumber space is carried out by the Fast Fourier Transfer (FFT), which makes the HOS method numerically efficient.

When provided with a frequency directional spectrum, $S(\omega, \theta)$, the discrete HOS mode amplitudes are initialized as

$$\frac{|B_{mn}^{\eta}|^{2}}{2\Delta k_{x}\Delta k_{y}} = S_{HOS}(k_{x}, k_{y}) = \frac{\partial \omega}{\partial k} \frac{1}{k} S(\omega, \theta)$$

$$B_{mn}^{\tilde{\phi}} = \frac{-\mathrm{i}g}{\omega} \beta_{mn}^{\eta}$$
(1)

where B_{mn}^{η} and $B_{mn}^{\tilde{\phi}}$ are, respectively, for the free surface displacement η and the surface velocity potential $\tilde{\phi} \equiv \phi|_{z=\eta(x,t)}$, and $g = 9.81 \text{ m/s}^2$ is the gravitational acceleration. A random phase is assumed for the complex amplitude B_{mn}^{η} . Note that $B_{mn}^{\tilde{\phi}}$ is initialized by relating to B_{mn}^{η} via the linearization (only at t = 0) of the dynamic free surface boundary condition (i.e. the Bernoulli equation at the free surface). Since B_{mn}^{η} is the mode amplitude in the wavenumber space (k_x, k_y) , it is directly related to the spectrum $S(k_x, k_y)$ by definition. The latter is so related to $S(\omega, \theta)$ that the total energy represented by the two is the same.

The public-domain HOS-ocean model presently offers the capability to simulate waves based on a directionally spread, JONSWAP-based spectrum with a single peak centered at $\theta = 0^{\circ}$ (i.e., waves entering domain perpendicular to left [x = 0] boundary). Users can modify the peak frequency, spectral steepness, and directional spread of this spectrum, but they cannot prescribe a different peak direction or multiple peaks. The first new capability described in this report adapts and extends the original JONSWAP option to allow the user to specify two different peaks (with individually defined energy levels, steepnesses, and spreads) that enter the model domain from different, user-specified peak directions. The second newly implemented capability fully enables the input of WW3-formatted spectra data. While this option apparently was listed in the current public-domain version of HOS-ocean (as of 2018), it is in fact erroneous. The main issue is caused by the 2D interpolation of the input spectral data onto the HOS-ocean model grids. In the 2018 release of the HOS-ocean model, the bilinear and bicubic interpolation are not properly coded. To address the issue, we install the library of bivariate interpolation of scattered data based on the quadratic Shepard method (Renka, 1988). Other adjustments have to apply (which, however, do not change the energy distribution described by the original spectral data), in order to make $S(\omega, \theta)$ consistent with the HOS-ocean formalism and suitable to be used in equation (1); see Section 2.2.

This paper details the modifications made to HOS-ocean in order to implement these new capabilities, and describes a number of tests conducted with the new software. The modifications themselves are outlined in Section 2. Following this, Section 3 introduces several basic test cases that were developed and run for each new capability. Further discussion and conclusions are provided in Section 4, and information on dissemination of this code is supplied in Section 5. Copies of the subroutines that are related to the implementation are included in the Appendix.

2 MODIFICATIONS TO HOS-OCEAN

The new software created for HOS-ocean was added by modifying existing files, rewriting or changing existing subroutines and adding new variables as needed. The list of FORTRAN90 source files has not been changed by these modifications. The primary file that was altered is initial_condition.f90, but minor additional modifications were also made to the files variables_3D.f90, input_HOS.f90, and HOS-ocean.f90. Some changes were also required in the model input file (input HOS.dat), including the addition of 13 new lines providing for

user specification of parameters for each of the two new options described below. A sample copy of the new input file is also included in the Appendix.

2.1 Bimodal Spectra Based on JONSWAP Parameters

To facilitate an additional 2D spectral capability, we added routines allowing the user to specify a directionally bimodal input spectrum, featuring energy peaks at two separate directions. With this option, two JONSWAP-format spectra with user-specified shape, peak frequency, peak direction, and directional spreading are merged into a single frequency-directional spectrum.

The primary subroutine that has been added to enable this functionality is *initiate_cross_sea* (in HOS-ocean file initial_condition.f90). This subroutine is constructed from modified sections of the existing subroutine *initiate_irreg*, in which initialization of single-peak 2D spectra based on JONSWAP is done. In the new subroutine, two separate copies of the JONSWAP-building lines are utilized (with minor modifications) to generate the two separate peaks describing the crossing sea state (Gramstad et al., 2018).

The processing of the input file (input_HOS.dat) in Fortran file input_HOS.f90 is modified to allow the user to specify a new case number (*icase* = 33) along with input parameters for each of the two desired peaks (see Appendix for sample input file). The input specifications for wave height (*Hs_real*), peak period (*Tp_real*) and JONSWAP parameter (*gamma_J*) are the same for the new case as for the single-peak option (*icase* = 3), except that there are separate sets of values for each of the two peaks (i.e., *Hsa*, *Tpa*, *gamma_a* and *Hsb*, *Tpb*, *gamma_b*). For this case, however, the directional spread of each peak is computed based on a cosine-power expression where the user specifies the power (i.e., N = Na and N = Nb) instead of the value *beta* that is used for single-peak spectra.

$$D(k_x, k_y) = \cos^N(\theta_{x, y} - \theta_p)$$
⁽²⁾

$$S_{HOS}(k_x, k_y) = \frac{C_g}{\left|\vec{k}\right|} S_{HOS}(k_x, k_y) \cdot D(k_x, k_y)$$
(3)

Finally, the user must specify the peak direction associated with each of the peaks (*theta_pa*, *theta_pb*) in degrees, corresponding to the principal directions toward which the two sets of waves are progressing when they enter the domain at the left ($k_x = 0$) boundary. It is recommended that these values remain between -90° and 90°.

2.2 Enabling Input of Frequency-Directional Spectra in WW3 Format

The public-domain HOS-ocean model is configured to read in a spectrum from single ASCII text file with a format similar to a standard WW3 output file. However, the model initialization in the public-domain version is not implemented correctly, as noted in Introduction. We have completely rewritten the subroutines *read_irreg_f* and *initiate_irreg_f* (both in the file initial_condition.f90)¹ to enable the simulation of waves using such fully 2D frequency-directional spectral data.

In the revised subroutine *read_irreg_f*, the following tasks are performed:

¹ The default input filename expected by the revised model is "WW3_spec_in.txt" (hard-coded in file variables 3D.f90).

(a) Read in the 2D spectrum from the input data file, assuming the WW3 format, and transform the data from the standard WW3 units of $m^2/(Hz rad)$ to the HOS-ocean units of $m^2/(rad/s rad)$, i.e.,

$$S(\omega,\theta) = \frac{S_{WW3}(f,\theta)}{2\pi}$$

(b) Preliminary calculations to determine the period T_p and direction of the maximum peak, and the total energy (which determines the energy-based significant wave height H_s). Note that HOSocean model runs in the non-dimensional space, and converts input quantities to dimensionless values before beginning the simulation. While the model input file (input_HOS.dat) requires the user to specify H_s and T_p for the normalization, those are over-written upon the preliminary calculations; i.e. in the case considered here, H_s and T_p given in input_HOS.dat are simply the placeholders to avoid run-time errors.

(c) Re-order the data so that the wave directions are in the ascending order. The sorting is necessary because the wave directions in the original WW3 data file are often not provided in a consistently increasing order, while those in HOS-ocean are ordered.

(d) Shift the axis of wave direction to re-locate the energy peaks in the range $-\pi/2 < \theta < \pi/2$. Currently, HOS-ocean model domain is limited to the upper half plane, and uses the above range for the wave directions. However, the WW3-formatted spectra are given in the range $0 < \theta < 2\pi$. A shift $\Delta\theta$ (user-specified) is applied to all wave angles, so that the major energy peaks of the input spectrum are contained in $-\pi/2 < \theta < \pi/2$. (Note that any waves from the original S_{WW3} that end up outside this angle range will be cut off and not represented in the HOS simulations. However, the energy lost due to the truncation is expected to be negligible since ocean waves are rarely so broad-banded.)

It should be mentioned that the procedures of re-ordering and shifting in (c) and (d) do not alter the energy distribution (i.e., the shape of the spectrum) described by the original input data. It is assumed that the user knows the overall distribution of energy densities in the original spectrum, and how much the spectral angles must be shifted in order to fit within the HOS-ocean domain as described above. (This may be determined, for example, simply by examining the input spectrum with a utility such as Matlab.) A new parameter (*ww3shft*) for specifying the spectral shift in degrees has been added to the input_HOS.dat file for use with *icase*=31. (This *icase* value was already reserved in HOS-ocean for WW3-formatted input files, but the option did not work.) The additional line is highlighted in the sample input file provided in the Appendix. This line must be included to avoid a run-time error; for an input spectral file that already has the wave energy bins properly situated for the HOS-ocean domain, the *ww3shft* parameter may be set to $0.0.^2$

At present, several diagnostic files are still written as output by *read_irreg_f* in the current version of HOS-ocean maintained at NRL, to allow the user to confirm that the data are correctly processed and shifted before beginning a simulation. These files include ww3_unsort_dimlss.dat (original unsorted data, but converted to units of m²/(rad/s rad) and then normalized), ww3 sort dimlss.dat (sorted and normalized), ww3 sorted dim.dat

² Note that, depending on the size that is defined for the HOS-ocean domain, some frequency bins of the input WW3 spectrum may also be excluded. This is dependent on parameters n1, n2, xlen, and ylen specified by the user as detailed on the website <u>https://github.com/LHEEA/HOS-ocean</u> and in Ducrozet et al. (2016), and it is beyond the scope of the present work.

(sorted, units of m²/(rad/s rad)), and ww3_rctr_dimlss.dat (recentered, units of m²/(rad/s rad)).³ Frequencies are normalized by the factor $(gk_p)^{-0.5}$, where g = 9.81 m/s² and k_p is the wavenumber corresponding the peak frequency ($f_p = 1/T_p$), while the spectrum is normalized by the factor $k_p^2 f_p$. Normalization and recentering should not change the shape of the spectrum. Consequently, when one plots these outputs in a utility such as Matlab, the graphs should look similar (other than the shifts in peak values and locations). Example plots of data from each of these files, initialized with a typical WW3 spectral file, are provided in Figure 1. The bottom two panels of the figure demonstrate how a user-specified shift of the spectrum can move peak energy closer to $\theta = 0$ and also allow a larger proportion of total spectral energy to be included in the HOS simulation.

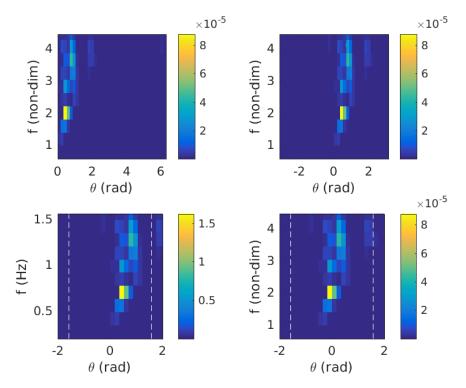


Figure 1. Spectral output of diagnostic test files generated at initiation of HOS-ocean simulation utilizing WW3 input spectrum that was shifted in direction using *ww3shft* = -30°. <u>Top left</u>: ww3_unsort_dimlss.dat; <u>Top right</u>: ww3_sort_dimlss.dat; <u>Bottom left</u>: ww3_sort_dimlss.dat; <u>Bottom right</u>: ww3_rctr_dimlss.dat. Dashed lines in bottom two panels indicate directional limits of HOS-ocean domain, i.e., $-\pi/2 < \theta < \pi/2$.

In the revised subroutine *initiate irreg f*, there are two tasks:

(a) The spectral data in *S*_{HOS} (which have been processed by *read_irreg_f*) are interpolated onto the grid-points of the HOS-ocean domain. Here, a bivariate (quadratic) interpolation library is called for the scattered data (Renka, 1988). The algorithm of interpolation is based on the Shepard method which is an inverse distance-weighted least squares method. To call the library,

³ Each of the diagnostic files is written in three-column format, with the columns containing frequency, direction, and spectral density respectively. Each column must be reshaped to a matrix of dimensions [# of directions] \times [# of frequencies].

a user needs to declare the following parameters: *n*, the number of data points; *nq*, the number of points to be used in least squares fitting; *nw*, the number of data points within the radii of influence (hence determining the radii); *nr*, dividing the surface containing the data into *nr*-by-*nr* cells (which optimizes the search of relevant data points). In the current implementation, the declaration is done in *initiate_irreg_f* (with the recommended values). Note that the interpolation must be done in the (ω, θ) plan, since the coordinate-transformation from (ω, θ) to (k_x, k_y) is nonlinear.

(b) The interpolated spectrum $S_{HOS,intr}(\omega, \theta)$ is then used to initialize the mode amplitudes; see equation (1). Note that the dimensionless form of equation (1) is coded. The complex mode amplitudes for the free surface displacement have random phase angles.

Following this, subroutine *initiate_irreg_f* returns control to the main program, and the computation of evolving wave states begins (utilizing existing subroutines *fourier_2_space* and *space_2_fourier*). From this point, the simulation proceeds in the same manner as in the public-domain model.

3 TEST CASES

Several test cases for evaluating the new capabilities with idealized input spectra are described in this section. First, both new model features are applied to the simulation of a single, directionally broad wave spectrum in Section 3.1. Then, in Section 3.2, the system is tasked with simulating waves generated from a directionally bimodal spectrum. For all cases, HOS-ocean is configured to utilize 256 modes in both x- and y-directions (i.e., n1 = n2 = 256), the wave peak period is fixed at 4.0 sec, and the wave domain is set to fit 45 individual peak waves in the x- and y-directions (i.e., xlen = ylen = 45 in the input_HOS.dat file, resulting in a domain that is approximately $1124 \times 1124 \text{ m}^2$). The generated wave state from the input spectrum (at t = 1 sec) is compared to that after a 100 sec simulation in each case, and the separate JONSWAP and WW3 cases are also compared. Possible sources of output differences between the two types of simulations are discussed.

3.1 Directionally Broad Single-Peak Spectrum

In this test set, two versions of an idealized spectrum with a single peak are first created with each of the two new input methods and configured to be nearly identical (Fig. 1), and then a third simulation is also run with a shifted version of the original spectrum. This case is primarily a test of the model's handling of a WW3-formatted input spectrum and how well the resulting simulation duplicates the wave state generated by the model's existing JONSWAP capability (which in this case did not require any code modifications). The boundary spectrum in both initial cases has a peak direction $\theta_p = 0^\circ$ (i.e., peak waves enter normal to the *y*-axis boundary with $k_y = 0$). The spectrum in the third case is shifted so that peak waves enter the domain at a 25° angle, providing a test of the model's consistency if such a shift should be required.

3.1.1 JONSWAP-based spectrum

This case was set up and run as a "control" case for comparison with the newly implemented WW3-formatted spectral input. The boundary spectrum in this case makes use of HOS-ocean's existing capability to produce a wave state based on a directionally spread spectrum. This capability is initialized in the model with parameters from a somewhat modified JONSWAP spectral formulation (Ducrozet et al., 2016):

$$S_{J}(\omega,\theta) = C_{J}H_{s}^{s}\omega_{*}^{2}\exp\left(-1.25\omega_{*}^{4}\right)\gamma^{\exp\left(-0.5(\omega_{*}-1)^{2}/\sigma^{2}\right)} \cdot (1/\beta)\cos^{2}\left(\frac{\pi\theta}{2\beta}\right)$$
(5)

where ω_* is wave frequency ω normalized by peak frequency ω_p , θ is wave direction in radians,

 γ is an adjustable parameter affecting the amplitude and narrowness of the spectral peak, and β is a directional spreading parameter developed by Dysthe (1979) with a functional range of roughly 0.10 (broad) to 0.30 (narrow). The (dimensionless) mode amplitudes for the free surface displacement coefficients of the wavenumber representation are computed from the spectrum as (Ducrozet et al., 2007)

$$\left|b_{k}\right| = \frac{g^{2}}{2\omega_{k}^{4}}S_{J}(\omega,\theta)\delta k_{x}\delta k_{y}$$

$$\tag{6}$$

assuming random phases for the complex b_k . For the test, the model input file was configured with the options shown in Table 1.

Parameter	Value
Duration of simulation (T_stop)	100 sec
Boundary wave format (icase)	3
Peak period (Tp_real)	4.0 sec
Peak direction	0° (default)
Sig wave height (Hs_real)	0.30 m
JONSWAP parameter (gamma_J)	3.3
Directionality (Dysthe beta)	0.30

Table 1. Input parameters for JONSWAP spectrum (single peak)

For this case, the spectral representation of the wave state was saved at the beginning and end of the 100 sec simulation. A comparison of the two spectra indicate that HOS-ocean did a good job of maintaining the prescribed wave state (Figure 2, top panels). The total variance of the input and output spectra differs by only 0.16%.

3.1.2 WW3 formatted input spectrum

This case was set up using the newly implemented, rewritten option for WW3 spectral input. A single-peak WW3 spectrum file was constructed, based directly on the initial JONSWAP file used in Section 3.1.1 (see Table 2). The spectrum in this input file was confirmed to have the same shape and nearly the same total variance (0.88% difference) as the spectrum used in the JONSWAP test.

Parameter	Value
Duration of simulation (T_stop)	100 sec
Boundary wave format (icase)	31
Peak period (Tp_real)	4.0 sec
Peak direction	0°
Sig wave height (from spectrum)	0.30 m

Table 2. Input parameters for WW3-formatted input spectrum (single peak)

The spectral representation of the wave state was again saved at the beginning (t = 1 sec) and end (t = 100 sec) of the simulation. A comparison of these two spectra indicates that HOS-ocean

again did a very good job of maintaining the prescribed wave state (Figure 2, middle panels). The variance of the initial and final spectra differs by just 0.03%.

While the difference in total variance between the JONSWAP-formatted and WW3-formatted input spectra is somewhat larger at 0.88%, as noted earlier, the output spectra differ by a comparable, slightly lower 0.67% (Figure 2, bottom panels). As seen in the figure, however, there is a noticeable difference between the distributions of the variance in the corresponding JONSWAP and WW3 spectra, which was not apparent in the comparison of these spectra at the time they were read in. This is evidenced by the *absolute* differences between JONSWAP and WW3 in/out spectra, which are much larger at 14% / 25%, respectively.

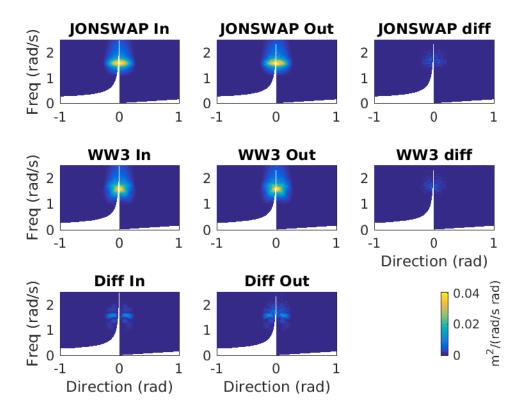


Figure 2. Comparison of initial and final spectral states, HOS-ocean simulation initialized with JONSWAP or WW3-formatted spectrum file. Difference spectra (right panels and bottom panels) show absolute differences on same scale as original spectra (lower right).

3.1.3 WW3 formatted input spectrum with direction shift

This case was set up using the same newly implemented WW3 spectral input format, with the spectral parameters shown in Table 3. A single-peak WW3 spectrum file was constructed in the same fashion as that in Section 3.1.2, except that now a shift of $+25^{\circ}$ was imposed on the spectrum before initiation of the simulation. In this case, the initial (t = 1 sec) spectrum had a somewhat narrower width and greater total variance (12% greater) in comparison to the spectrum used in the preceding section.

Parameter	Value
Duration of simulation (<i>T_stop</i>)	100 sec
Boundary wave format (<i>icase</i>)	31
Peak period (<i>Tp_real</i>)	4.0 sec
Original peak direction	0°
Shift in wave direction (<i>ww3shft</i>)	+25.0°
Sig wave height (from spectrum)	0.32 m

Table 3. Input parameters for *shifted* WW3-formatted input spectrum (single peak)

The spectral representation of the wave state was once again saved at the beginning and end of the simulation. A comparison of the two spectra indicate that, despite the generally skewed diagonal path of waves through the domain, HOS-ocean again did a very good job of

maintaining the prescribed wave state (Figure 3). The variance of the initial (t = 1 sec) and final (t = 100 sec) spectra in this case differs by just 0.13%.

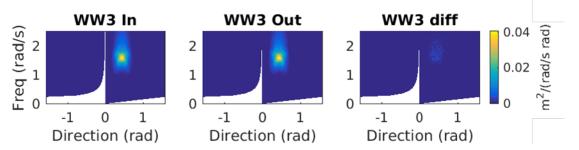


Figure 3. Beginning (*left*) and ending (*middle*) spectral states for 100 sec simulation with HOSocean, initialized by single peak spectrum shifted by $ww3shft = +25^{\circ}$. Difference between beginning and end states is shown in right panel.

3.2 Directionally Bimodal (Two-Peak) Spectrum

In this set of tests, we evaluate the newly implemented capability to generate two-peak spectra based on JONSWAP parameters with an idealized test case, and a comparative simulation is run using a directionally bimodal WW3 input spectrum created from the original bimodal JONSWAP input. The same system settings (i.e., n1, n2, xlen, ylen) and simulation period (i.e., $T_stop = 100$ sec) are used as in Section 3.1, but in this case the two peaks are each configured to be directionally narrow (cosine power N = 100).

3.2.1 JONSWAP-based spectrum

This case was set up using the newly implemented bimodal JONSWAP spectral input format. Specifying the new "icase" value of 33 in the HOS-ocean input file, we also prescribed JONSWAP parameter values for each of two separate spectral peaks (labeled "a" and "b", as shown in Table 4). In this case, the JONSWAP peakedness parameter was set to a very large value (*gamma* = 10.0) for each peak, and the cosine-power directional spreading was set to a very narrow value (N = 100).

Parameter	Value
Duration of simulation (T_stop)	100 sec
Boundary wave format (icase)	33
Peak period (Tpa, Tpb)	10 sec
Peak direction A (thet_pa)	+15°
Peak direction B (thet_pb)	-15°
Sig wave heights (Hsa, Hsb)	0.3 m
Sig wave height (total)	0.42 m
JONSWAP parameter (gamma_a,b)	10.0
Directional spread (Na, Nb)	100

Table 4. Input	parameters for	directionally	bimodal JONSWAI	P spectrum
----------------	----------------	---------------	-----------------	-------------------

Input and output spectra for the JONSWAP bimodal case are shown along with their absolute difference in the top three panels of Figure 4. A comparison of the two spectra indicate that HOS-ocean again did a reasonably good job of maintaining the prescribed wave state. The variance of the input and output spectra in this case differs by only 0.37%. Comparing the peaks

of this spectrum to the one generated with the parameters of Section 3.1.1, we find that these appear to be steeper (as expected from the larger *gamma* value) and to have a roughly similar directional spread. The similarity in spreads is not immediately anticipated, as two different methods ("Dysthe β " vs. "cosine power") vs were used to prescribe them. However, the result is at least somewhat foreseeable, as both the use of a value of *beta* = 0.30 in Section 3.1.1 and a cosine power N = 100 in this case are expected to produce directionally narrow spectral peaks.

3.2.2 WW3 formatted bimodal input spectrum

This case was set up using the newly implemented WW3 spectral input format (Table 5). A twopeak WW3 spectrum file was constructed, based directly on the initial directionally bimodal JONSWAP file used in Section 3.2.1. This input file was again confirmed to have the same shape and nearly the same total variance as the spectrum used in the bimodal JONSWAP test (Figure 4).

Parameter	Value
Duration of simulation (<i>T_stop</i>)	100 sec
Boundary wave format (<i>icase</i>)	31
Shift in wave direction (<i>ww3shft</i>)	0.0°
Peak directions (from spectrum)	-15°, 15°
Sig wave height (from spectrum)	0.43 m

Table 5. Input parameters for bimodal WW3-formatted spectrum

Spectra for the idealized WW3 bimodal case are shown at input (t = 1 sec) and output (t = 100 sec) time steps along with their absolute difference in Figure 4, middle panels. A comparison of the two spectra indicate that HOS-ocean maintained the prescribed wave state as well as in the preceding cases. The total variance of the input and output WW3 spectra in this case differs by just 0.02%. However, the absolute difference in variance was 11%, suggesting that the variation in phases and wave-wave interactions between the two directionally bimodal configurations caused the randomly phased waves to evolve in different ways that significantly shifted spectral energy levels during the simulation, although once again total energy was simultaneously conserved.

As in the single-peak tests, the initial (t = 1 sec) WW3 spectrum modeled on the JONSWAP spectrum in this case is very close to to the initial (t = 1 sec) JONSWAP spectrum. Total variance between WW3 and JONSWAP initial spectra differs by only 0.2%, and the absolute variance difference is just 0.6%. At the conclusion of the simulations (t = 100 sec), total variance of the WW3 and JONSWAP final spectra differs by less than 0.02%, but the absolute variance difference jumps to almost 19%. This may again reflect the differing pathways for nonlinear energy transfer among the randomly phased waves in the two separate simulations. While total energy in the domain is again well conserved in both simulations, interactions among different waves at different times seem to have produced considerable variations in the final spectral distributions (Figure 4, bottom panels).

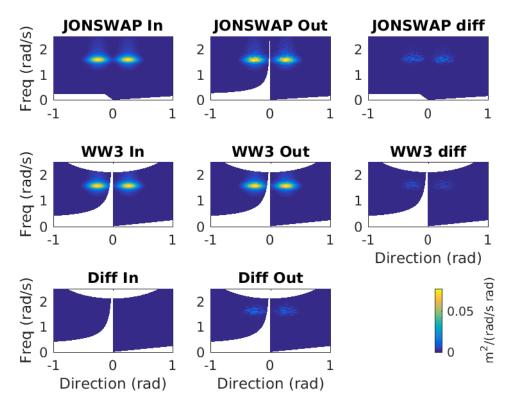


Figure 4. Comparison of initial and final spectral states, directionally bimodal HOS-ocean simulation initialized with JONSWAP or WW3-formatted spectrum file. Difference spectra (*right panels and bottom panels*) show absolute differences on same scale as original spectra (*lower right*).

3.2.3 WW3 formatted bimodal input spectrum with direction shift

This case was again set up using the newly implemented WW3 spectral input format, with parameters summarized in Table 6. A double-peak WW3 spectrum file was constructed in the same fashion as that in Section 3.2.2, except that now a shift of -75° was imposed on the spectrum before initiation of the simulation. Because of the extreme shift, a portion of one of the bimodal peaks was no longer in the angle range required for HOS-ocean (i.e., $\theta_{p,shift} = -15^\circ - 75^\circ = -90^\circ$ or $-\pi/2$). The total variance in the shifted spectrum (Figure 5) is now 5.8% less than that of the original spectrum.

Parameter	Value
Duration of simulation (<i>T_stop</i>)	100 sec
Boundary wave format (<i>icase</i>)	31
Shift in wave direction (<i>ww3shft</i>)	-75.0°
Initial peak directions (from spectrum)	-15°, 15°
Sig wave height (from spectrum)	0.41 m

Table 6.	Input	parameters	for sl	hifted	bimodal	WW3-	formatted s	pectrum

As with previous tests, the spectral representation of the wave state was again saved at the beginning (t = 1 sec) and end (t = 100 sec) of the simulation. A comparison of the two spectra indicate that HOS-ocean once again did a very good job of maintaining the prescribed wave state, based on the shifted spectrum which was partially cut off. The variances of the initial and

final spectra differ by 0.03%. Once again, the absolute difference is significantly larger, at 13%, indicating that energy levels were conserved but redistributed somewhat among wave frequencies and directions.

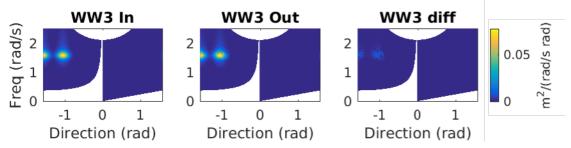


Figure 5. Beginning (t = 1 sec, *left*) and ending (t = 100 sec, *middle*) spectral states for simulation with HOS-ocean, initialized by originally symmetric bimodal spectrum shifted by *ww3shft* = -75°. Difference between beginning and end states is shown in right panel.

4 DISCUSSION AND CONCLUSIONS

We have implemented and tested two additional input options in the model HOS-ocean, which enable simulations to be initialized either with bimodal JONSWAP-based wave spectra or with more general frequency directional distributions in a WAVEWATCH III (WW3)-formatted input file. The bimodal JONSWAP-format spectra are generated using an adapted version of HOSocean's existing capability for creating directionally spread, single peak JONSWAP boundary spectra, with cosine power spreading utilized in the new capability instead of "Dysthe β ". The WW3-formatted spectral input and initialization were implemented by re-writing the subroutines that properly process and interpolate the input data, in accordance with the HOS-ocean requirements. Both capabilities have been validated with simple test cases, producing the results as anticipated.

For initializing single-peak spectra, it was confirmed that simulations with the modified HOSocean were substantially identical, whether they were initialized with JONSWAP parameters or with a WW3-formatted spectrum file based on those parameters. Total variance values for the initial and final spectra were nearly the same between the two sets of simulations, although energy was redistributed among the frequency directional bins to some degree during the simulations. When a directional shift was applied to the single-peak WW3 spectrum (from a peak direction of 0° to a new peak of 25°), this resulted in a 12% increase in overall spectral variance, possibly because the model was unable to fully capture (in wavenumber space) the spectral energy of waves traveling near angle $\theta = 0^\circ$ for the specific domain that was used. This could probably be remedied by specifying larger values of domain parameters *n1, n2, xlen*, and *ylen*.

For initializing double-peak spectra (i.e., energy peaks at two different directions), it was confirmed that simulations with the modified HOS-ocean were again substantially identical, whether they were initialized with JONSWAP parameters or with a JONSWAP-based, WW3-format spectrum file. Variance values of the initial spectra were again nearly the same between the two sets of simulations, and in this case the variances of final spectra again differed significantly (19%) between the JONSWAP and WW3-initialized cases. This may have been due to differences in the random phases of the two cases and their effects on nonlinear interactions among crossing waves from the bimodal wave state. In a third case, a large directional shift was applied to the double-peak WW3 spectrum (from a mean direction of 0° to a

new mean of -75°), this resulted in a 5.8% decrease in overall spectral variance, primarily because a portion of one bimodal peak was shifted outside of the HOS-ocean directional range (i.e., $-90^{\circ} < \theta < 90^{\circ}$) so that some if its waves were excluded from the resulting simulations. Despite the shift, HOS-ocean still maintained the overall variance of the waves throughout the simulation.

Additional tests are presently being conducted with this code to investigate the effects of bimodal wave spectra on the development of extreme (rogue) ocean waves. These extended investigations will be completed in early FY21 and will be detailed in Part II of this memorandum report series.

5 DISSEMINATION OF CODE

HOS-ocean is an open-source, "community" model, available in its public release version from the GitHub platform (<u>https://github.com/LHEEA/HOS-ocean</u>). The new features described herein will ultimately be included in a future public release version of the software, but at present the full set of modifications may be obtained by contacting one of the report authors.

6 REFERENCES

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7 APPENDIX – PRINTOUTS OF NEW CODE AND INPUT FILE

As noted earlier, the new code added to HOS-ocean was primarily contained in subroutines *initiate_cross_sea, read_irreg_f*, and *initiate_irreg_f*, all of which are included in the HOS-ocean file initial_condition.f90. Each of these new subroutines is printed out below. An example of a modified input file, input HOS.dat, is also included following the code printouts.

```
SUBROUTINE initiate cross sea
I _____
                           _____
! Written by Jie Yu, NRL, 2018-08-21
! -Initialisation of crossing sea state (case=33) that consists of more than
! one wave systems such that the wave spectrum has multiple peaks with
! different directions and frequencies
! -In this version, the crossing sea state is a random bimodal wave field
  (Gramstad et al 2018), given by S(ome,dir)=Sa(ome,dir)+Sb(ome,dir), where
!
 Sa abd Sb are two JONSWAP spectra with their own directional distributions
!
  (cos(thet-thet p))**N.
1
1
! -The parameters specifying the spectra,
   Tpa, thet pa, Hsa, gamma a, Na, and Tpb, thet pb, Hsb, gamma b, Nb
!
! are given in input HOS.dat
1
! -In this version, deep-water waves are assumed.
|-----
1
   implicit none
   integer :: iseed, i1, i2
   real(rp) :: pioxlen, pioylen
   real(rp) :: angle, angle1, angle2
   real(rp) :: alp_a, alp_b, sigma, coeff_a, coeff_b
   real(rp) :: a, ome dim, omega pa, omega pb
   real(rp) :: E a, E b, E tot, Cg
   real(rp), dimension(mlo2p1,m2,2) :: rnd
   real(rp), dimension(mlo2p1,m2) :: Sa, Sb, Da, Db, Sc
   real(rp), dimension(m1o2p1,m2) :: thet hos
   integer, dimension(2)
                                :: indx
   pioxlen = TWOPI/xlen star !dkx
   pioylen = TWOPI/ylen star
                              !dky
   omega pa= TWOPI/Tpa
   omega_pb= TWOPI/Tpb
   thet pa = thet pa*PI/180.0 rp !change from deg to rad
   thet pb = thet pb*PI/180.0 rp
! Directional spreading functions
   !Cal theta used in HOS. By definition, thet hos = [-pi, pi]
   thet hos(1,1) = 0.0 rp
   do i1=1,n1o2p1
     do i2=1,n2
       if (i1 /=1 .or. i2 /= 1) then
         thet_hos(i1,i2) = atan2(ky_n2(i2), kx(i1))
       endif
     enddo
   enddo
```

```
!coefficients An in D(theta)
   coeff a = gamma(Na/2.0 rp+1.0 rp)/gamma(Na/2.0 rp+0.5 rp)/sqrt(PI)
   coeff b = gamma(Nb/2.0 rp+1.0 rp)/gamma(Nb/2.0 rp+0.5 rp)/sqrt(PI)
   !Da(theta), Db(theta)
   do i1=1,n1o2p1
     do i2=1,n2
       if (abs(thet hos(i1,i2)-thet pa) .le. PIO2) then
          Da(i1,i2)=coeff a*(cos(thet hos(i1,i2)-thet pa))**Na
       else
          Da(i1,i2)=0.0 rp
        endif
       if (abs(thet hos(i1,i2)-thet pb) .le. PIO2) then
         Db(i1,i2)=coeff b*(cos(thet hos(i1,i2)-thet pb))**Nb
       else
         Db(i1,i2)=0.0 rp
       endif
      enddo
   enddo
! Sa(omega) and Sb(omega)
! T = 1/omega_pa is used for normalization
   Sa(1,1)=0.0 rp
   Sb(1,1)=0.0 rp
   do i1=1, n1o2p1
     do i2=1,n2
       if (i1 /=1 .or. i2 /= 1) then
         ome dim=omega n2(i1,i2)/T
         if (ome dim .lt. omega pa) then
            sigma=0.07 rp
          else
           sigma=0.09 rp
          endif
          a=exp(-(ome dim-omega pa)**2/(2.0 rp*sigma**2*omega pa**2))
         Sa(i1,i2)=omega_n2(i1,i2)**(-5)* &
            exp(-5.0 rp/4.0 rp*(ome dim/omega pa)**(-4))*gamma a**a
          if (ome dim .lt. omega pb) then
            sigma=0.07_rp
          else
            sigma=0.09 rp
         endif
         a=exp(-(ome dim-omega pb)**2/(2.0 rp*sigma**2*omega pb**2))
          Sb(i1,i2)=omega n2(i1,i2)**(-5)* &
            exp(-5.0 rp/4.0 rp*(ome dim/omega pb)**(-4))*gamma b**a
       endif
      enddo
   enddo
   if (iseven(n2)) then
      Sa(:, n2o2p1) = 0.0 rp
     Sb(:, n2o2p1) = 0.0 rp
   endif
   !Determine alp_a and alp_b, s.t. E_a = (Hsa/4) **2, E_b = (Hsb/4) **2
   E a = 0.0 rp
   E b = 0.0 rp
   do i1 = 1, n102p1
     do i2 = 1, n2
       if (i1 /= 1 .or. i2 /= 1) then
         Cg = 0.5_rp*omega_n2(i1,i2)/k_abs(i1,i2) !deep-water Cg
         E_a = E_a + Cg/k_abs(i1,i2)*Sa(i1,i2)*Da(i1,i2)*pioxlen*pioylen
         E_b = E_b + Cg/k_abs(i1,i2)*Sb(i1,i2)*Db(i1,i2)*pioxlen*pioylen
       endif
      enddo
```

```
enddo
   alp a = (Hsa/4.0 rp)**2/(grav**2*omega pa**(-4))/E a
   alp b = (Hsb/4.0 rp)**2/(grav**2*omega pa**(-4))/E b
   !Recall that omega pa is used for normalization
   write(*,*) 'In initiate_cross_sea(): alp_a, alp_b = ', alp_a, alp_b
   Sa = Sa*alp a
   Sb = Sb*alpb
! Amplitudes of the Fourier modes
   !random numbers used for the phases
   if (random phases.eq.0) then
     !same random numbers for each run, given (n1,n2)
     call init not random seed()
     call random number(rnd)
   elseif (random phases.eq.1) then
     !different random numbers for each run
     call init random seed()
     call random number(rnd)
   else
     print*, 'Random number generation undefined'
     stop
   endif
   Sc
       = 0.0 rp
   a eta = 0.0 cp
                     !make sure a_eta(1,1)=0, a_phis(1,1)=0
   a phis= 0.0 cp
   E tot = 0.0 rp
   do i1 = 1, n102p1
     do i2 = 1, n2
       if (i1 /= 1 .or. i2 /= 1) then
         Cg = 0.5 rp*omega_n2(i1,i2)/k_abs(i1,i2)
                                                      !deep-water Cg
         Sc(i1,i2) = Sa(i1,i2) * Da(i1,i2) + Sb(i1,i2) * Db(i1,i2)
         angle1 = rnd(i1,i2,1)*TWOPI
         angle2 = rnd(i1,i2,2)*TWOPI
         angle = 0.0_rp
         a eta(i1,i2)=(2.0 rp*Cg/k abs(i1,i2)*Sc(i1,i2)*pioxlen*pioylen)**(0.5 rp) &
                       *exp(i*(angle1+angle2+angle))
         a_phis(i1,i2)=(-1.0_rp*i/omega_n2(i1,i2))*a_eta(i1,i2)
         E_tot=E_tot + 0.5_rp*abs(a_eta(i1,i2))**2
       endif
     enddo
   enddo
   E tot = E tot * g star ! g star=1 for deep-water wave
   indx=maxloc(Sc)
   il=indx(1)
   i2=indx(2)
   write(*,*) 'In initiate cross sea,g star,L out,T out=',g star,L out,T out
   write(*,*) 'For Sc=Sa+Sb, Tp =',TWOPI/(omega n2(i1,i2)/T),' thet p = ', &
                thet hos(i1,i2) *180.0 rp/PI
   write (*,*) 'E tot =',E tot, 'Hs ini =',4*SQRT(E tot/g star) * L out
   open(99,file='cross_sea_spc.dat',status='unknown')
   do i1=1, n1o2p1
     do i2=1,n2
      write(99,*) omega n2(i1,i2)/T,thet hos(i1,i2), &
       Sc(i1,i2)*grav**2*omega_pa**(-5), Sa(i1,i2)*grav**2*omega_pa**(-5), &
       Sb(i1,i2)*grav**2*omega pa**(-5), Da(i1,i2), Db(i1,i2)
     enddo
   enddo
   close(99)
```

SUBROUTINE read irreg f 1 !-----! Rewritten by Jie Yu, NRL. 2018-08-08 - Read in the spectrum given by a WAVEWATCH III simulation ! - Recenter the spectrum along the direction-axis, so that the main peaks ! are within the HOS angle range [-pi/2, pi/2] 1 ! - Calculate Tp real and Hs real ! - In this version, deep-water waves are assumed. ! Note: If the input spectrum is already properly situated in [-pi/2,pi/2], skip the steps of shifting and relocation. 1 !-----1 IMPLICIT NONE REAL(RP) :: dthet, O p ww, k p ww, E int tot, variance INTEGER :: i1,i2,ii,jj,ind_o_max !real(rp) :: Tp_real, L_out, T_out, HS_real !REAL(RP), DIMENSION(ifreq) :: phi_ww_int,E_int,df !REAL(RP), DIMENSION(7) :: phi temp !real(rp), dimension(ifreq,ithet) :: phi2 !real(rp), dimension(ithet) :: thet2 !integer, dimension(ithet) :: indx !integer, dimension(2) :: temp !** Move the commented declarations to variables 3d.f90 !REAL(RP), ALLOCATABLE :: freq ww(:) !REAL(RP), ALLOCATABLE :: thet_ww(:) !REAL(RP), ALLOCATABLE :: phi_www(:,:) !REAL(RP), ALLOCATABLE :: phi_ww_int,E_int,df REAL(RP), DIMENSION(7) :: phi_temp !real(rp), ALLOCATABLE :: phi2(:,:) !real(rp), ALLOCATABLE :: thet2(:) !integer, ALLOCATABLE :: indx(:) integer, dimension(2) :: temp integer nloc character*31 str1, str2 integer :: ind real(rp) :: shft,tmp1,tmp2 $! \ {\rm shft} \ {\rm is} \ {\rm for} \ {\rm relocating} \ {\rm the} \ {\rm spectral} \ {\rm energy}. \ {\rm Optimal} \ {\rm value} \ {\rm may} \ {\rm be}$! determined by visual inspection of the WW3 data (e.g., using matlab). ! User specifies it as input ww3shft in input HOS.dat file. shft = ww3shft*PI/180.0 rp ! Shift value specified by user (convert to radians) ! Read WW3 spectrum data file open (1002,file=ww3 filename,status='old') read(1002,*) str1,ifreq,ithet,nloc,str2 npt = ifreq*ithet nr = floor(dsqrt(npt/3.0 rp)) allocate(freq_ww(ifreq), thet_ww(ithet), phi ww(ifreq, ithet)) allocate(phi ww int(ifreq), E int(ifreq), df(ifreq)) allocate(phi2(ifreq,ithet)) allocate(thet2(ithet), indx(ithet))

```
DO ii=1, FLOOR(ifreq/8.0)
       READ(1002,*) freq_ww(8*(ii-1)+1:8*ii)
      ENDDO
      if (8*FLOOR(ifreg/8.0).lt.ifreg) then
       ii=8*floor(ifreq/8.0)+1
       read(1002,*) freq ww(ii:ifreq)
      endif
      DO ii=1, FLOOR (ithet/7.0)
       READ(1002,*) thet ww(7*(ii-1)+1:7*ii)
      ENDDO
      if (7*FLOOR(ithet/7.0).lt.ithet) then
       ii=7*floor(ithet/7.0)+1
       read(1002,*) thet ww(ii:ithet)
      endif
      dthet=abs(thet ww(2)-thet ww(1))
      !df as define in plot_ww3_2dspec.m
      df(1)=freq ww(2)-freq ww(1);
      df(ifreq) = freq ww(ifreq) - freq ww(ifreq-1);
      do ii=2, ifreq-1
       df(ii)=(freq ww(ii+1)-freq ww(ii-1))/2.0 rp;
      enddo
      READ(1002,*)
      READ(1002,*)
      DO i1=1, FLOOR(ithet * ifreq / 7.0)
       READ(1002,*) phi temp(1:7)
       DO i2 = 1,7
          jj=FLOOR(((i1-1)*7+i2-1)/REAL(ifreq,RP))+1
          ii=(i1-1)*7+i2 - (jj-1)*ifreq
         phi_ww(ii,jj)=phi_temp(i2)
                                           !phi ww(freq,dir)
        ENDDO
      ENDDO
      READ(1002,*) phi ww(ifreq-(ithet *ifreq-7*FLOOR(ithet*ifreq/7.0))+1:ifreq,ithet)
      CLOSE (1002)
! Preliminary calculations
      temp=MAXLOC(phi ww)
      write(*,*) 'In read irreg f, T p =',1.0/freq ww(temp(1)), &
                 ' peak dir =', 360.0/TWOPI*thet ww(temp(2))
      E int
               = 0.0 rp
      E_{int_tot} = 0.0_{rp}
      variance = 0.0 rp
      DO jj=1,ithet
       DO ii = 1, ifreq
         E int(ii) = E int(ii) + phi ww(ii,jj) * dthet
         variance = variance + phi ww(ii,jj) * dthet * df(ii)
       ENDDO
      ENDDO
      DO ii = 1, ifreq-1
        E int tot = E int tot + \&
           0.5 rp*(E int(ii)+E int(ii+1))*(freq ww(ii+1)-freq ww(ii))
      ENDDO
      write(*,*) 'In read irreg f, 4sqrt(variance) =',4.0 rp*sqrt(variance)
      write(*,*) 'In read_irreg_f, 4sqrt(E_int_tot) =',4.0_rp*sqrt(E_int_tot)
! Transform frequency to omega, and S(f,dir) to S(omega,dir)
```

I.

```
freq ww=freq ww * TWOPI
                                   !omega=2*pi*freq
      phi ww=phi ww / TWOPI
                                    !S(omega)=S(f)/2pi
| ********
! Adim the variables (preliminary)
| **********
     DO ii=1, ifreq
       phi ww int(ii)=SUM(phi ww(ii,:)) * dthet
     ENDDO
! Locate the peak freq
     ind o max= MAXLOC (phi ww int, 1)
     O p ww = freq ww(ind o max)
     k p ww = 0 p ww^{*2} / grav
                                       !assuming deep-water waves
     Tp real = TWOPI / O_p_ww
     L out
              = 1/k p ww
              = 1/0_p_ww
      Tout
     Hs real = 4.0 rp * sqrt(E int tot)
              = sqrt(grav*k_p_ww)
     tmp1
              = k_p_ww **2 * 0_p_ww
     tmp2
     write(*,*) 'From WW3 file, Hs real, Tp real =', Hs real, Tp real
     write(*,*) 'From WW3 file, o p ww, sqrt(grav*k p ww),k p ww^2*o p ww =',&
        O_p_ww, tmp1, tmp2
!In the original WW3 file, the wave directions are not ordered ('messed up')
!Sort phi ww in ascending angle
      open(77,file='ww3 unsort dimlss.dat',status='unknown')
      open(78,file='ww3 sort dimlss.dat',status='unknown')
      open(79,file='ww3_sort_dim.dat',status='unknown')
     open(80, file='ww3 rctr dimlss.dat', status='unknown')
     ind=minloc(thet ww,1)
     thet2(1)=thet ww(ind)
     indx(1) = ind
      do ii=2,ithet
       ind=minloc(thet_ww,1,mask=thet_ww.gt.thet2(ii-1))
       thet2(ii)=thet_ww(ind)
       indx(ii) =ind
      enddo
      do ii=1, ifreq
       do jj=1,ithet
         phi2(ii,jj) = phi ww(ii,indx(jj))
       enddo
      enddo
      do ii=1,ifreq
       do jj=1,ithet
         write(77,*) freq ww(ii)/tmp1, thet ww(jj), phi ww(ii,jj)*tmp2
         write(78,*) freq ww(ii)/tmp1, thet2(jj), phi2(ii,jj)*tmp2
         write(79,*) freq ww(ii),
                                     thet2(jj),
                                                     phi2(ii,jj)
       enddo
      enddo
     close(77)
     close(78)
     close(79)
     phi ww=phi2
     thet ww=thet2
     temp=maxloc(phi ww)
     !Sorting should not change the peak freq and dir
     write(*,*) 'After sorting, T p =',TWOPI/freq ww(temp(1)), &
                          ' peak dir =',360.0/TWOPI*thet ww(temp(2))
! Relocate the peaks
     do ii=1,ithet
       thet_ww(ii) = thet_ww(ii)+shft
       if(thet_ww(ii) .ge. TWOPI) thet_ww(ii) = thet_ww(ii)-TWOPI
      enddo
      ind=minloc(thet ww,1)
```

```
thet2(1)=thet ww(ind)
     indx(1) = ind
     do ii=2,ithet
       ind=minloc(thet ww,1,mask=thet ww.gt.thet2(ii-1))
       thet2(ii)=thet ww(ind)
       indx(ii) =ind
     enddo
     do ii=1, ifreq
       do jj=1,ithet
         phi2(ii,jj) = phi ww(ii,indx(jj))
       enddo
     enddo
     phi ww=phi2
     thet ww=thet2
     temp=maxloc(phi ww)
     !Tp real should not be changed
     write(*,*) 'After recentering, T p =',TWOPI/freq ww(temp(1)), &
                      ' peak dir =',360.0/TWOPI*thet ww(temp(2))
! Admin variables (after sorting and relocation)
     !Check the peaks and total energy
     do ii=1, ifreq
       phi ww int(ii)=SUM(phi ww(ii,:)) * dthet
     enddo
     !Locate the peak freq
     ind o max= MAXLOC(phi ww int,1)
     O p ww = freq ww(ind o max)
     k_p ww = 0_p w**2 / grav
                                     !deep water
     Tp real = TWOPI / O_p_ww
     L out = 1/k p ww
     T out = 1/0 p ww
     Hs real = 4.0 \text{ rp} + \text{sqrt}(\text{E int tot})
     tmp1 = sqrt(grav*k_p_ww)
tmp2 = k_p_ww **2 * 0_p_ww
     write(*,*) 'After relocation, Hs_real, Tp_real =', Hs_real, Tp_real
     do ii=1,ifreq
       do jj=1,ithet
         write(80,*) freq ww(ii)/tmp1,thet ww(jj), phi ww(ii,jj)*tmp2
       enddo
     enddo
     close(80)
100 continue
     !Normalization. Prepare for initiate irreg f
     freq_ww = freq_ww/tmp1
     phi_ww = phi_ww*tmp2
END SUBROUTINE read irreg f
_____
SUBROUTINE initiate irreg f
1
|-----
! Rewritten by Jie Yu, NRL, 2018-08-08
! Initialisation of irregular multi-directional sea state (linear)
  - From spectrum file from WAVEWATCH III
1
   - Bivariate interpolation onto model grid points.
! Note:
! Make sure that the WW3 spectrum peaks are in the range [-pi/2, pi/2],
  which is currently the HOS setup. The modification is done in read irreg f().
1
! It can also be pre-processed using matlab.
```

```
!-----
1
IMPLICIT NONE
INTEGER :: iseed, i1, i2
REAL(RP) :: theta, E, pioxlen, pioylen
REAL(RP) :: angle, angle1, angle2
REAL(RP), DIMENSION(m1o2p1,m2,2) :: rnd
REAL(RP), DIMENSION(mlo2p1,m2) :: phi_E
INTEGER, DIMENSION(mlo2p1,m2) :: ind_o_ww,ind_t_ww
! For interpolation
! n = number of data points
Т
  ng= number of points to be used in least square fit.
       5 <nq< min(40, n-1). nq=13 is highly recommended.
!
1
  nw= number of data points within (hence determining) the radii of influence R(k).
!
    1 <nw< min(40,n-1). nw=19 is recommended for large data set
1
  nr: dividing the surface containing the data into nr x nr cells.
!
      nr = sqrt(n/3) is recommended.
1
  !* Move to variables 3d: integer (kind=4), parameter :: npt= ifreq*ithet
  integer (kind=4), parameter :: ng = 13
  !* Move to variables 3d: integer(kind=4), parameter :: nr = floor(dsqrt(npt/3.0 rp))
  integer (kind=4), parameter :: nw = 19
  integer (kind=4), dimension(nr,nr) :: lcell
  integer (kind=4), dimension(npt) :: lnext
  real (kind=8), dimension(npt)
                                   :: rsq
  real (kind=8),
                  dimension(npt) :: xd, yd, zd
  real (kind=8), dimension(5,npt) :: coeff
 real (kind=8), dimension(nlo2p1,n2) :: thet_hos
  integer (kind=4) :: ier
  real (kind=8)
                :: px,py,q,dx,dy,xmin,ymin,rmax
                  :: qx,qy !needed if qs2grd() is used
  real (kind=8)
  real (kind=8)
                 :: phi max, cutoff
  integer (kind=4) :: temp(2)
  pioxlen = TWOPI/xlen star
                              !dkx
  pioylen = TWOPI/ylen star
                              !dky
  do i1=1, ifreq
   do i2=1, ithet
     xd((i1-1)*ithet + i2) = freq ww(i1)
      yd((i1-1)*ithet + i2) = thet ww(i2)
      zd((i1-1)*ithet + i2) = phi_ww(i1,i2)
    enddo
  enddo
! Rescale to make zd max=1.0, to avoid small numbers in interpolation cal.
  phi max = maxval(phi ww)
 zd = zd/phi max
  cutoff = 0.05 rp
! Call QSHEP2 to define the interpolant Q to this data.
1
  call qshep2(npt,xd,yd,zd,nq,nw,nr,lcell,lnext,xmin,ymin,&
             dx,dy,rmax,rsq,coeff,ier)
  if (ier /= 0) then
   write (*,'(a,i8)') ' Error in qshep2, ier = ', ier
   stop
  end if
```

```
!Calculate theta used in HOS
thet hos(1,1) = 0.0 rp
i2=1
do i1=2,n1o2p1
 thet hos(i1,i2) = atan2(ky n2(i2), kx(i1))
enddo
i1=1
do i2=2,n2
 thet hos(i1,i2) = atan2(ky n2(i2), kx(i1))
enddo
do i1=2,n1o2p1
 do i2=2,n2
   thet hos(i1,i2) = atan2(ky n2(i2), kx(i1))
  enddo
enddo
phi E = 0.0 rp
do i1=1, n1o2p1
 do i2=1,n2
    px=omega n2(i1,i2)
    if(px .ge. freq_ww(1) .and. px .le. freq_ww(ifreq)) then
      py=thet_hos(i1,i2)
      if (py .ge. thet_ww(1) .and. py .le. thet_ww(ithet)) then
        !Do the interpolation.
        q=qs2val(px,py,npt,xd,yd,zd,nr,lcell,lnext, &
                 xmin, ymin, dx, dy, rmax, rsq, coeff)
        !If gradient at (px,py) is also needed use qs2grd
        !!call qs2grd(px,py,npt,xd,yd,zd,nr,lcell,lnext,xmin, &
        !!
                      ymin, dx, dy, rmax, rsq, coeff, q, qx, qy, ier)
        if (q .lt. 0.0) then
          if (abs(q) .le. cutoff) then
            q = 0.0 rp
          else
            write(*,'(a,3e14.6)') 'At (px,qx), negative q : ', px,py,q
            !stop
          endif
        endif
        phi E(i1, i2) = q
      else
       phi_E(i1,i2) = 0.0_rp !Outside the range of directions
      endif
    else
      phi E(i1, i2) = 0.0 rp
                               !Outside the range of frequencies
    endif
  enddo
enddo
if (iseven(n2)) phi E(:,n2o2p1) = 0.0 rp
phi E = phi E*phi max !scale back
write(*,*) 'max phi_E =', maxval(phi_E)
write(*,*) 'n1o2p1, n2 = ', n1o2p1, n2
open(99,file='ww3_intpl.dat',status='unknown')
do i1=1,n1o2p1
 do i2=1,n2
    px=omega n2(i1,i2)
   py=thet hos(i1,i2)
    write(99,*) px,py,phi E(i1,i2)
 enddo
enddo
close(99)
```

! Compute the random numbers used for phases

```
IF (random phases.EQ.0) THEN
    ! Same random numbers for each run, given (n1,n2)
    CALL init not random seed()
    CALL RANDOM NUMBER (rnd)
  ELSEIF (random_phases.EQ.1) THEN
    ! Different random numbers for each run
    CALL init random seed()
    CALL RANDOM NUMBER(rnd)
  ELSE
    PRINT*, 'Random number generation undefined'
    STOP
  ENDIF
! -Initialize the amplitudes of the Fourier modes for surface elevation \eta
! and surface velocity potential.
     a eta = 2*(Cg/k)*S(omega, theta)*dkx*dky
!
    a phis= (-i*g/omega)*a eta (linearized Beroulli eq at z=0, with p atm=0)
1
! -In the following implementation, deep-water waves are assumed.
! -Under the normalization used by HOS-ocean,
     2*(Cg/k)=1/omega*^3, -i*g/omega = -i/omega*,
!
! where omega* = omega/omega p is the dimensionless frequency.
  a eta=0.0 cp
                   !make sure a eta(1,1)=0, a phis(1,1)=0
  a phis=0.0 cp
  E=0.0 rp
 do i1 = 1, n1o2p1
    do i2 = 1, n2
      if (phi E(i1,i2) .gt. tiny) then
        !phi E should be non-negative after the interpolation calculations.
        !for small phi E, a eta = 0
        if (i1 /= 1 .or. i2 /= 1) then
          angle1 = rnd(i1,i2,1)*TWOPI
          angle2 = rnd(i1,i2,2)*TWOPI
          angle = 0.0 \text{ rp}
          a_eta(i1,i2)=(phi_E(i1,i2)*pioxlen*pioylen)**(0.5_rp) &
                        *(1.0 rp/omega n2(i1,i2)**3)**(0.5 rp) &
                        *exp(i*(angle1+angle2+angle))
          a phis(i1,i2)=(-1.0 rp*i/omega n2(i1,i2))*a eta(i1,i2)
          E=E + 0.5 rp*abs(a eta(i1,i2))**2
        endif
      endif
    enddo
  enddo
  E_tot = E * g_star ! g_star=1 for deep-water wave
 write(*,*) 'In initiate irreg_f, g_star,L_out,T_out=',g_star,L_out,T_out !jyu
write(*,*) 'E =',E,'Hs_ini =',4*SQRT(E/g_star) * L_out, 'Tp_ini =',Tp_real
END SUBROUTINE initiate irreg f
```

Modified input HOS.dat (new options highlighted in **bold blue** font):

Restart previous computation	:: i_restart	::	0
Choice of computed case	:: i_case	::	33
Geometry of the horizonta	al domain		
Length in x-direction	:: xlen	::	70.0
Length in y-direction	:: ylen	::	18.0

Time stuff (dimensionless Duration of the simulation Sampling frequency (output)	::			7200.0 2.0
Tolerance of RK scheme	::	toler	::	1.0e-7
Dommermuth initialisation	::	n	::	4
Dommermuth initialisation	::	Та	::	10.0
Physical dimensional parameters				
Gravity	::	grav	::	9.81
Water depth	::	depth	::	-50.
Irregular waves (i case=3)				
Peak period in s	::	Tp_real		10.0
Significant wave height in m	::	Hs real	::	1.4142
JONSWAP Spectrum	::	gamma_J	::	1.0
JONSWAP Spectrum Directionality (Dysthe)	::	beta	::	0.302
Random phases generation	::	random phases	::	1
WW3 input spectrum (i_case=31)				
Shift of WW3 spectrum deg	::	ww3shft	::	30.0
Crossing sea (i_case=33)				
Peak period in s	::	Тра	::	4.0
Peak direction deg	::	thet_pa	::	15.0
Significant wave height in m	::	Hsa		0.3
JONSWAP Spectrum	::	gamma_a	::	10.0
Direction spreading		Na	::	100
Peak period in s	::	Tpb	::	4.0
Peak direction deg	::	thet_pb	::	-15.0
Significant wave height in m	::	Hsb	::	0.3
	::	gamma_b	::	10.0
Direction spreading	::	Nb	::	100
Output files				
		1	::	11
Output: 1-dim. ; 0-nondim.	::	i_out_dim	::	1
3d free surface quantities	::	i_3d	::	1
3d modes		i_a_3d	::	0
2d free surface, center line			::	
Wave probes in domain			::	0
Swense output 1='yes',0='no'	::	i_sw	::	0
